PDS/SBN Rosetta Review VIRTIS

S. Protopapa 6-11 October 2017

Instrument

The VIRTIS (Visible Infra Red Thermal Imaging Spectrometer) instrument combines a double capability: (1) high-resolution visible and infrared imaging in the 0.25-5 micron range at moderate spectral resolution (VIRTIS-M channel) and (2) high-resolution spectroscopy in the 2-5 micron range (VIRTIS-H channel). The two channels will observe the same comet areas in combined modes to take full advantage of their complementarities.

VIRTIS-M (named -M in the following) is characterized by a single optical head consisting of a Shafer telescope combined with an Offner imaging spectrometer and by two two-dimensional FPAs: the VIS (0.25-1 micron) and IR (1-5 micron). VIRTIS-H (-H) is a high-resolution infrared cross-dispersed spectrometer using a prism and a grating. The 2-5 micron spectrum is dispersed in 8 overlapping orders on a focal plane detector array.

Test Case/VIRTIS-M

Detection of exposed H2O ice on the nucleus of comet 67P/Churyumov-Gerasimenko as observed by Rosetta OSIRIS and VIRTIS instruments

Barucci et al. 2016 A&A 595, A102 (2016)

N.	Observation	Cube	Start time	End time	Phase	Incidence	Emission	Δ	LST	T
	name	Parameters	(UT)		(deg)	(deg)	(deg)	(km)	(h)	(K)
1	I1_00383518966	432, 256, 158	2015-02-25T21:04:00	2015-02-25T21:30:19	53.02	66.94	41.92	81.51	12.23	203
2	I1_00385906923	432, 256, 70	2015-03-25T12:23:18	2015-03-25T12:46:34	73.49	67.08	45.83	88.23	15.20	197
3	I1_00385885107	432, 256, 70	2015-03-25T06:19:42	2015-03-25T06:42:58	74.47	53.03	31.27	94.06	12.43	218
4	I1_00373462192	432, 256, 86	2014-11-01T11:31:03	2014-11-01T11:45:22	103.07	60.11	43.28	32.39	11.04	168
5	I1_00377182711	432, 256, 80	2014-12-14T12:59:43	2014-12-14T13:13:02	91.77	43.41	52.15	19.44	15.44	188
6	I1_00376302211	432, 256, 80	2014-12-04T08:24:43	2014-12-04T08:38:01	91.07	57.34	70.64	23.49	14.81	179
7	I1_00369356914	432, 256, 109	2014-09-14T23:09:43	2014-09-14T23:45:57	66.89	78.22	30.26	28.16	11.03	163
8	I1_00377184571	432, 256, 74	2014-12-14T13:30:43	2014-12-14T13:43:02	92.66	48.30	54.87	19.92	15.59	158

Notes. For each spot, we report the observations offering the best signal-to-noise conditions with the pixel position (sample and line) reported in Table 4. For each pixel, basic information about observation time, geometry conditions, distance between the Rosetta spacecraft and comet surface (Δ), Local Solar Time (LST), and retrieved temperature (T) are given. The integration time is 3 s for all data reported. The cube parameters indicate the size of the acquisition in bands, sample, line dimensions.

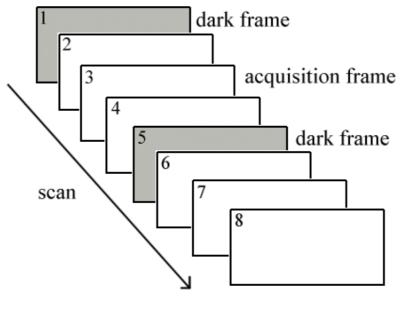
stp045/i1_00383518966

data storage for -M channel ROSETTA - VIRTIS EAICD

Document No.	: VIR-INAF-IC-001
ssue/Rev. No.	: issue 4.4
Date	: November 2015
Page	: 38

Raw file .qub INT= Array[432, 256, 158]Cal file .cal FLOAT= Array[432, 256, 150]

Raw data form the Visible and the IR FPA are stored in different files as Qube objects, so as to simplify data handling. For each FPA, science data and dark frames are stored in the same file, interleaved in the Qube core, and appear in the order they are transmitted (Fig. 4.2). Dark frames are acquired at intervals given by the DARK_ACQUISITION_RATE parameter and are temporally stored in the QUBE along the scan.



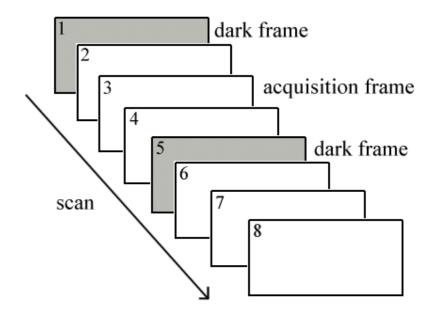
DARK_ACQUISITION_RATE=3

Fig. 4.2 – Structure of dark frames acquisition

The QUBE contains the data calibrated in radiance. Dark frames are removed from the qube, which is therefore smaller than the original raw data qube along the Z/time axis.

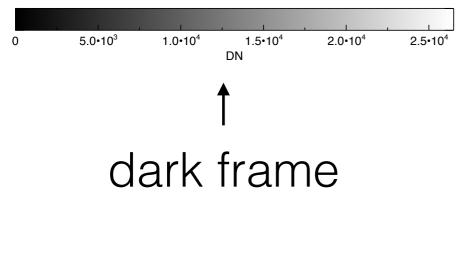
data storage for -M channel



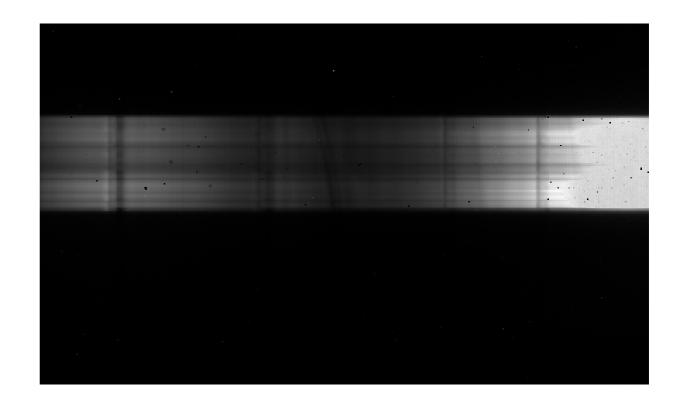


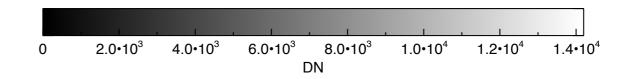
DARK_ACQUISITION_RATE=3

Fig. 4.2 – Structure of dark frames acquisition



acquisition frame \rightarrow





data storage for -M channel

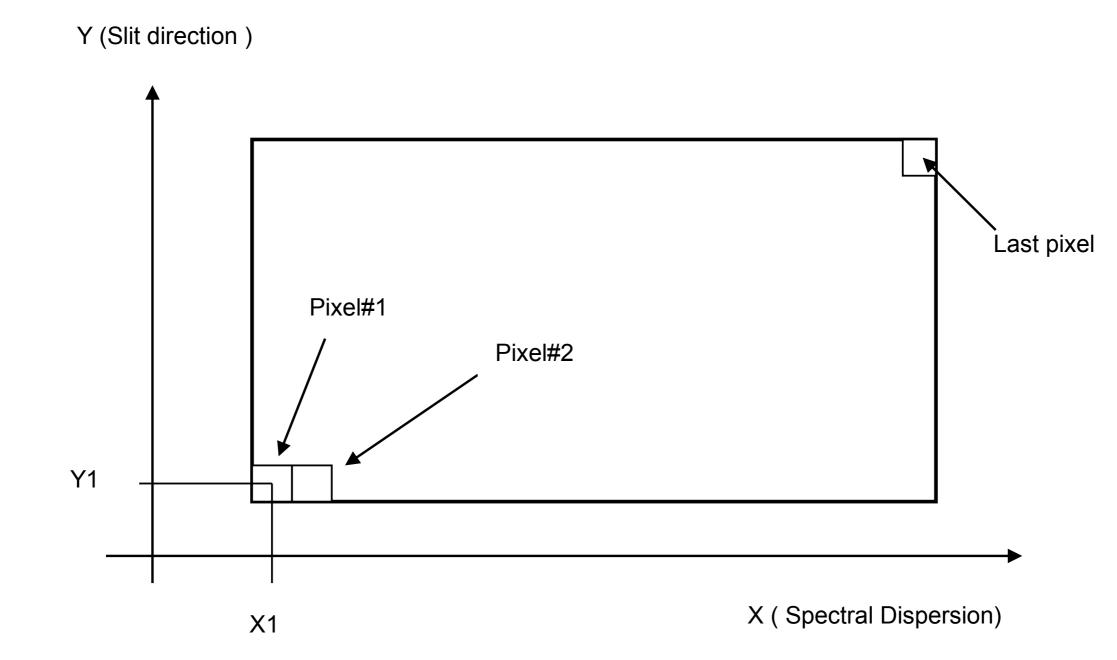


Figure 2.2 Data Organization on the CCD and IR detectors of VIRTIS-M

readpds .cal file

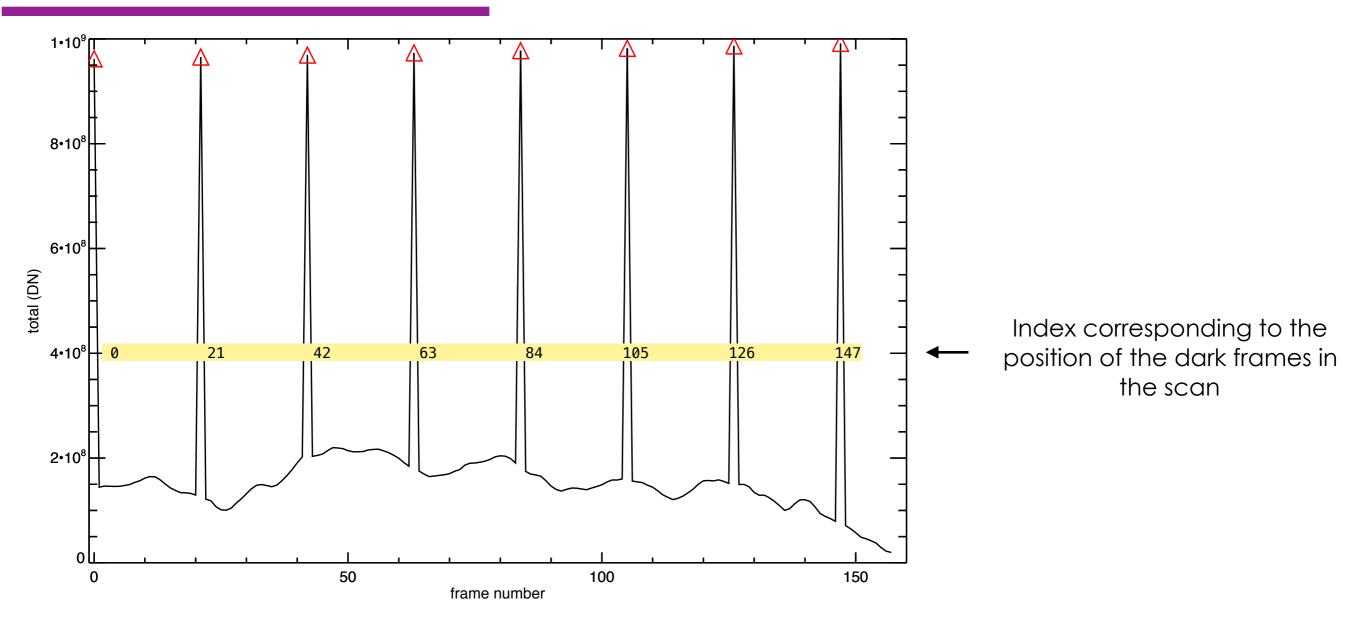
5-Oct-2017 2	22:27:24		Printed By: S	Silvia F	Protopapa	Sheet A
	ESA	Review It	em Discrepancy	,	Rosetta EOM Panel:	Rosetta_EOM_ Science_Panel
(A) TITLE:	readpds .cal file warning		RID No:	Scie	nce_Panel-265	
			ID No:	2134	13	
Orig. Ref.:	VIRTIS-US-SP-016		Originator:	Silvia	a Protopapa	
(B) DATAPA	CK DOCUMENT:					
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readpds .qub file

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		ID No:	213	344		
Orig. Ref.:	VIRTIS-US-SP-017	Originato	r: Silv	ria Protopapa		
(B) DATAPA	CK DOCUMENT:					
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Page/Sectio	n/Para:					
(C) DISCREI	PANCY:					
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Description	of Discrepancy:					
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(E) PANEL F	RECOMMENDATIONS:					
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1. step/reading the data

I selected the dark frames in the cube empirically, looking at the total DN per frame. My findings are consistent with the dark_acquisition_rate parameter

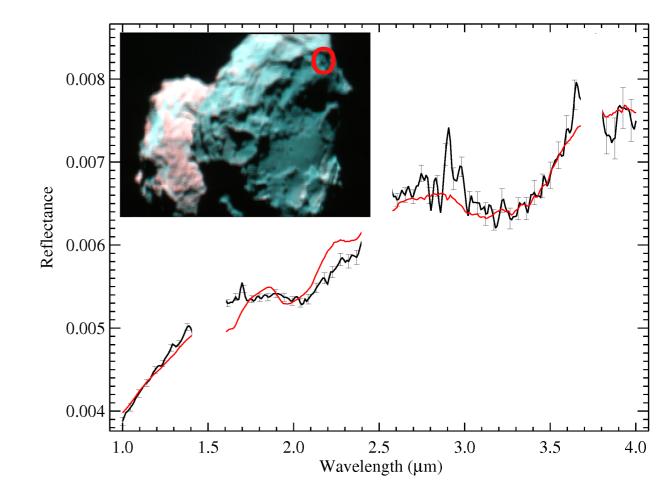


FRAME_PARAMETER = (3, 1, 10, 20) FRAME_PARAMETER_DESC = ("EXPOSURE_DURATION", "FRAME_SUMMING", "EXTERNAL_REPETITION_TIME", "DARK_ACQUISITION_RATE")

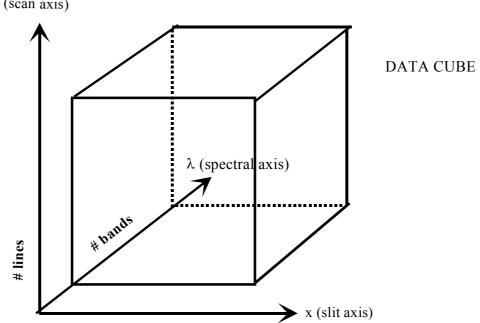
y (scan axis) **67P** λ (spectral axis) 7 # bands # lines

S. Protopapa

1000 2000 3000 4000 0 DN



Barucci et al. 2016 A&A 595, A102 (2016)



samples

Calibration pipeline file analyzed = stp045/i1_00383518966 Strange Pattern

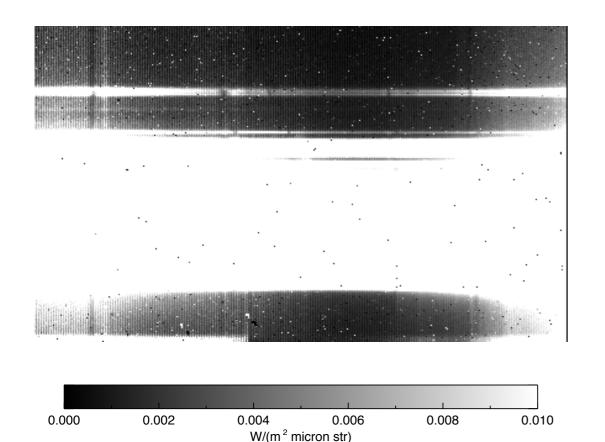
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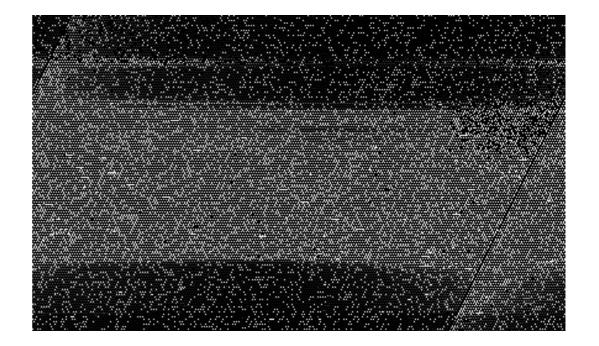
 $t_{IR} = 3sec$

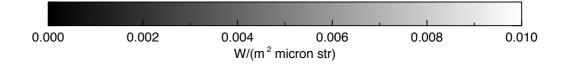
Considering the exposure time and the responsivity of the instrument, I convert the data from DN to RAD (see left panel)

S. Protopapa

data delivered to PDS







Response function

6-Oct-2017 1	7:36:16		Printed By:	Silvia F	Protopapa	S	heet A
	ESA	Review Item	Discrepan	су	Rosetta EOM Panel:	Rosetta_ Science	
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	correct		ID No:	2133	35		
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(B) DATAPA	CK DOCUMENT:						
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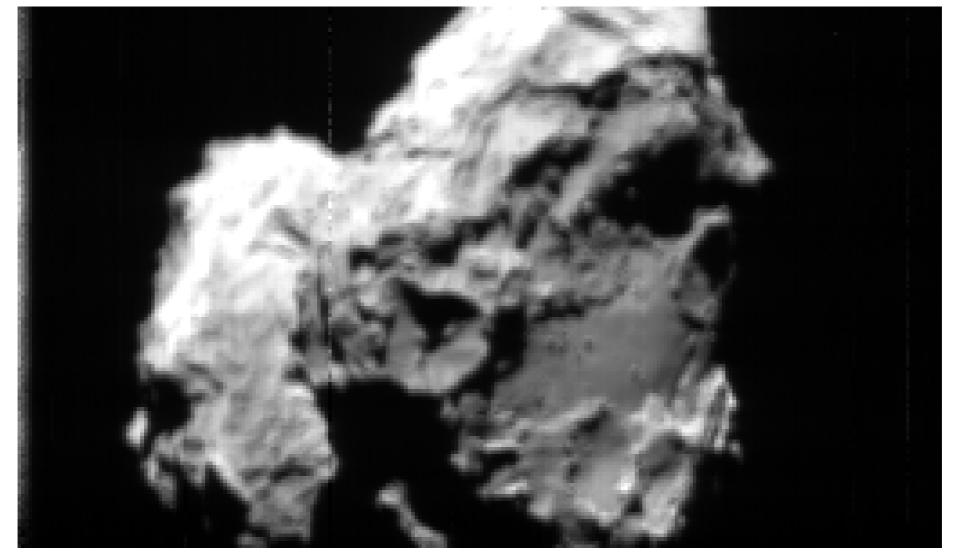
Wavelength calibration 6-Oct-2017 17

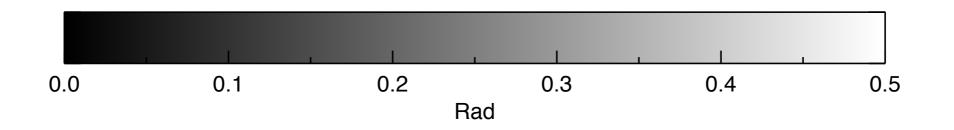
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						Rosetta EOM	Rosetta_E		
		ESA	Revie	w Item Discrepancy	y	Panel:	Science_Pa		
	(A) TITLE:	VIRTIS-M spectral calib	ration	RID No:	Scier	nce_Panel-259			
		table/documentation		ID No:	2133	2			
	Orig. Ref.: VIRTIS-US-SP-012 Originator: Silvia Protopapa								
	(B) DATAPACK DOCUMENT:								
	RO-SGS-PR-1018-App20: VIRTIS Rosetta EOM Science Archive Review Procedure (Issue:1.0) ROSETTA – VIRTIS EAICD/page 20/paragraph 2.4.5								
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Calibration pipeline

I considered the data I calibrated in RAD and generated a spectral cube #samples X #lines X #bands

Band = 10 (idl convention)

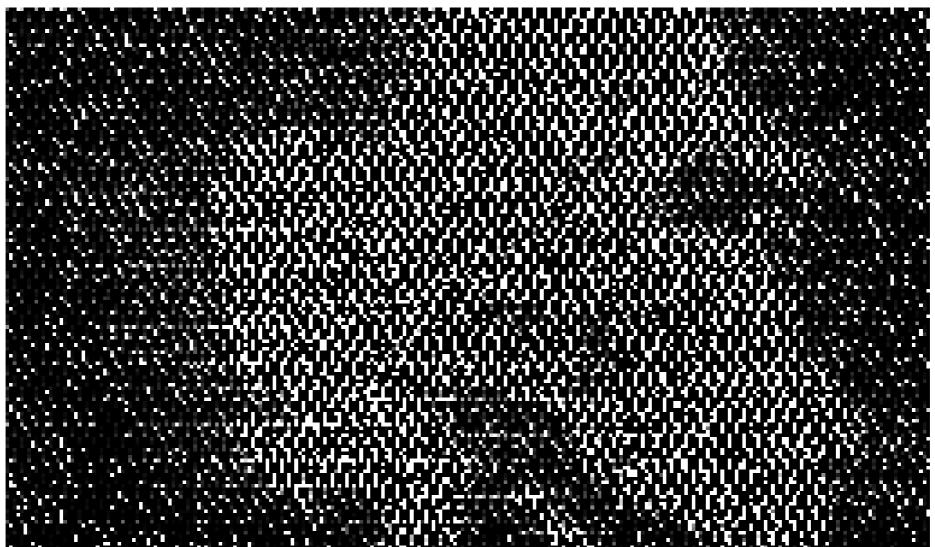


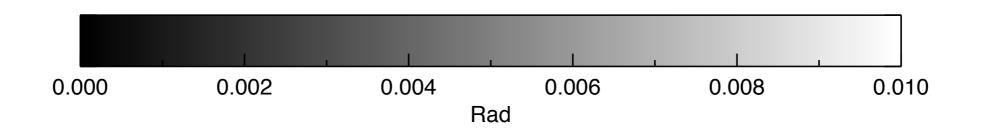


Calibration pipeline

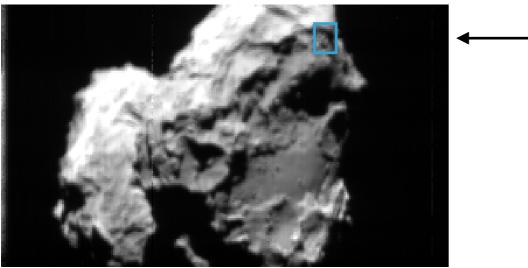
Data cube generated starting from the calibrated data provided to PDS. The pattern is funny. The orientation is different

Band = 10 (idl convention)



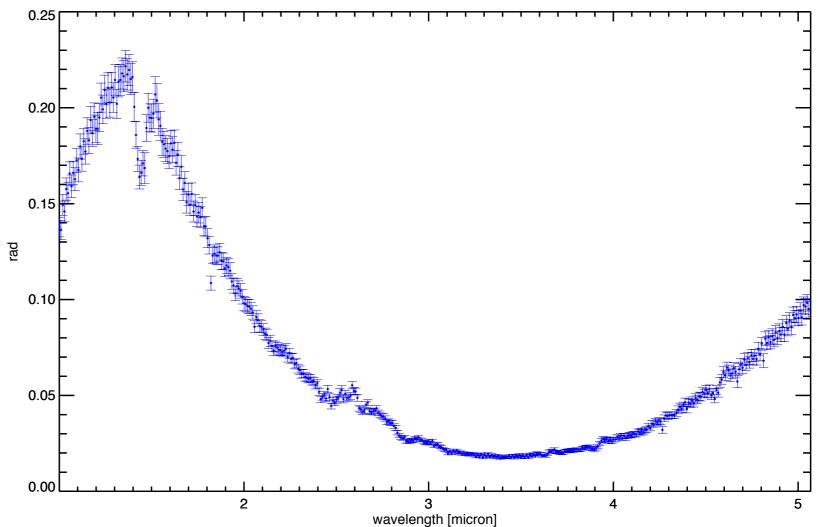


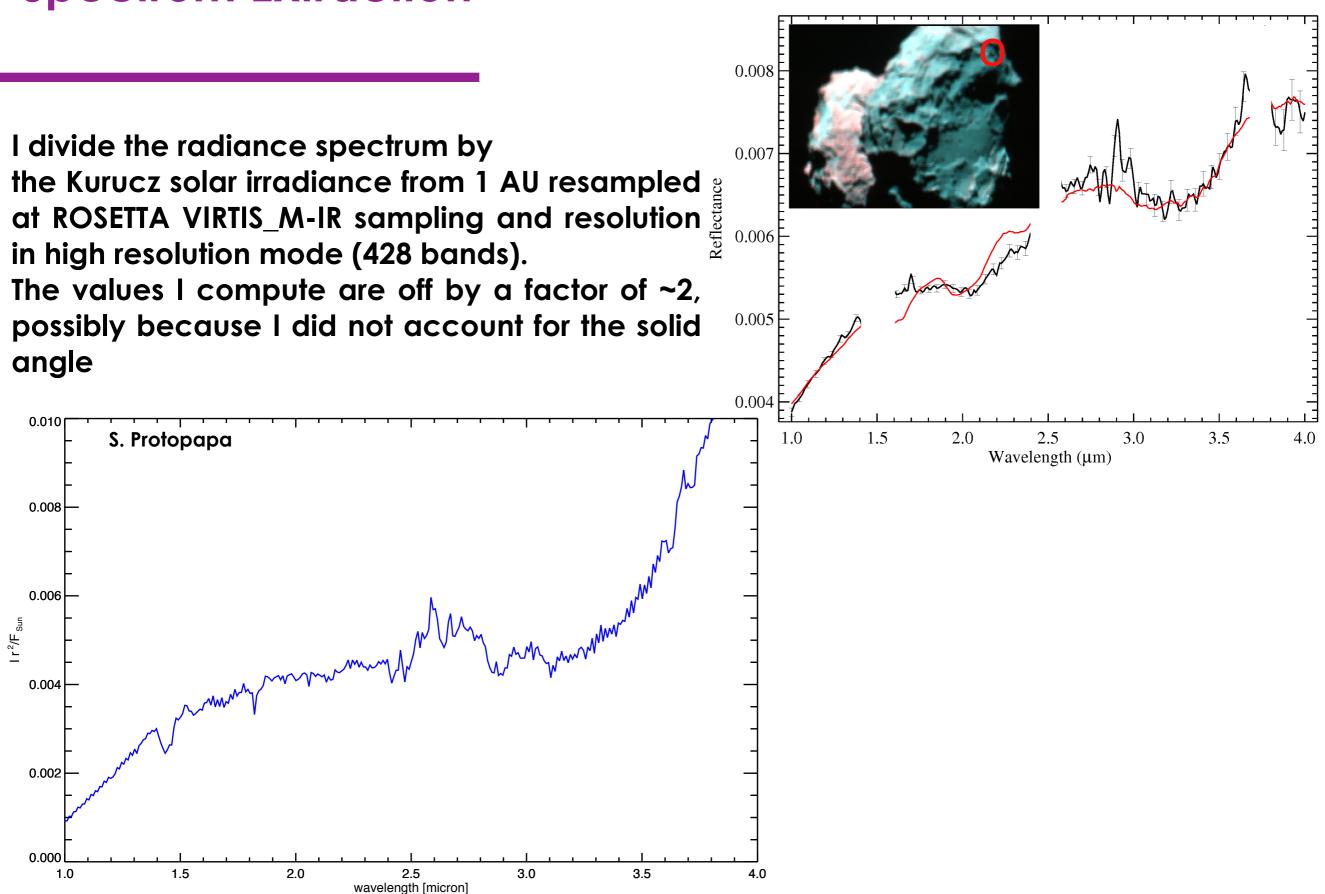
Spectrum Extraction



		1			1
0.0	0.1	0.2	0.3	3 0.4	0.5
		•			
			Rad		

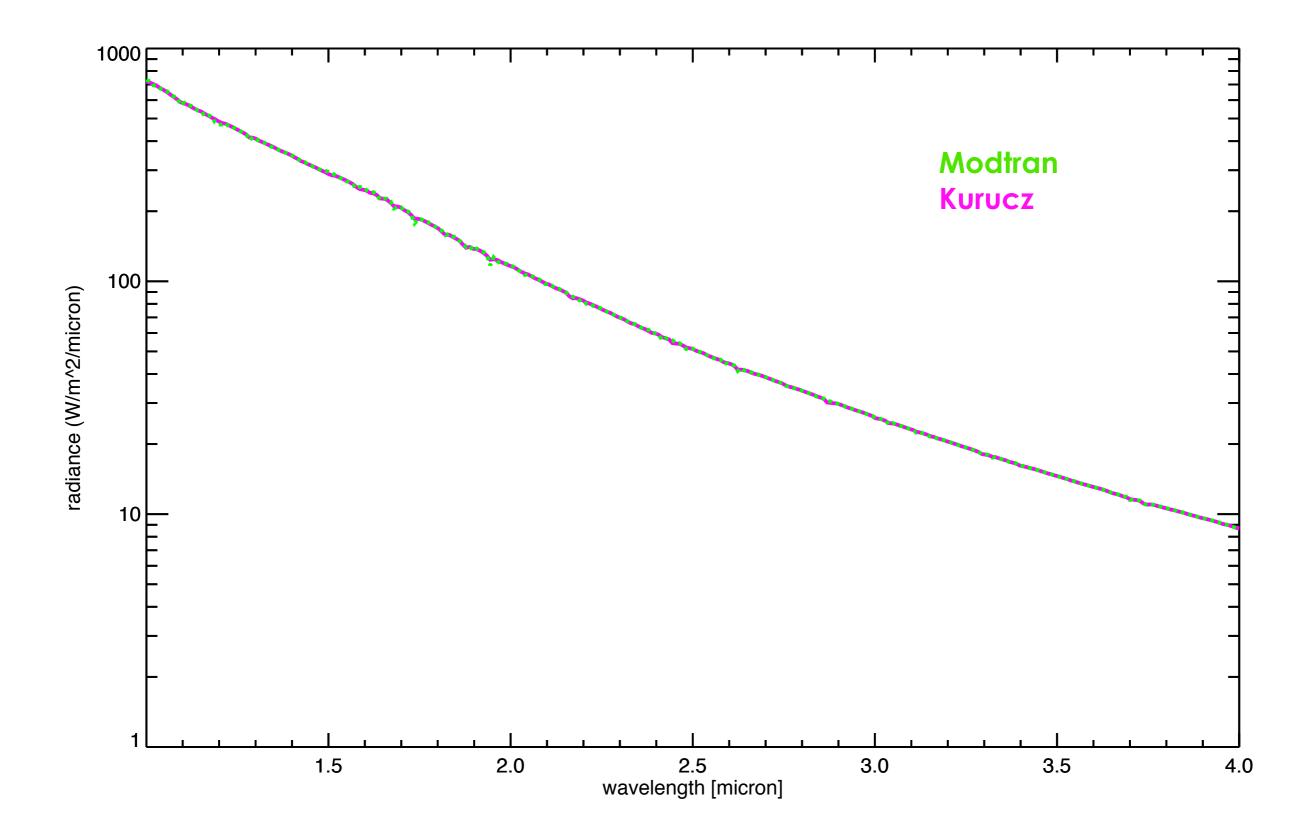
A sigma clipping algorithm is used to identify outliers. The value at each wavelength is the mean of the good pixels in the box and the uncertainty is given by rms deviation of the good pixels.





Spectrum Extraction

Calibration pipeline



Kurucz solar irradiance

5-Oct-2017 2	22:37:34		Pri	nted By:	Silvia	Protopapa	She
	ESA	Revie	w Item D	iscrepar	су	Rosetta EOM Panel:	Rosetta_EO Science_Pa
(A) TITLE:	virtis_res_ir_high_v10.lbl	/number of	rows F	RID No:	Scie	ence_Panel-268	
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Bad pixel map

5-Oct-2017 2	22:45:41		Printed By:	Silvia P	rotopapa	S	heet A
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Flat-fielding

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	ESA	Review Item Discrepanc	y Rosetta E Panel:	—
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VIRTIS-H

5-Oct-2017 2	22:25:05		Printed By:	Silvia I	Protopapa	Sheet /
	ESA	Review Ite	em Discrepan	су	Rosetta EOM Panel:	Rosetta_EOM_ Science_Panel
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